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# Fennoscandia before Nuna/Columbia: Paleomagnetism of 1.98–1.96 Ga mafic rocks of the Karelian craton and paleogeographic implications



N.V. Lubnina<sup>a,\*</sup>, S.A. Pisarevsky<sup>b,g</sup>, A.V. Stepanova<sup>c,d</sup>, S.V. Bogdanova<sup>e,f</sup>, S.J. Sokolov<sup>c</sup>

<sup>a</sup> Faculty of Geology, M.V. Lomonosov Moscow State University, Leninskiye Gory, 1, Moscow 119991, Russia

<sup>b</sup> Earth Dynamics Research Group, ARC Centre of Excellence for Core to Crust Fluid Systems (CCFS) and The Institute for Geoscience Research (TIGeR), Department of Applied Geology, Curtin University, GPO Box U1987, WA 6845, Australia

<sup>c</sup> Institute of Geology, Karelian Research Center, Pushkinskaya st., 11, Petrozavodsk 185910, Russia

<sup>d</sup> Institute of Geology of Ore Deposits, Petrography, Mineralogy and Geochemistry RAS, Staromonetny, 35, Moscow 119017, Russia

<sup>e</sup> Department of Geology, Lund University, Sölvegatan 12, SE 223 62 Lund, Sweden

<sup>f</sup> Kazan Federal University, 18 Kremlyovskaya str., Kazan 420008, Russia

<sup>g</sup> Institute of the Earth's Crust, Siberian Branch of the Russian Academy of Sciences, Irkutsk, 128 Lermontov Street, 664033, Russia

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## ABSTRACT

Numerous mafic dykes, sills and intrusions with ages between 1985 Ma and 1960 Ma are exposed near the Onega Lake in southern Karelia, Russia. The paleomagnetic analysis of these rocks has revealed a stable remanence with directions belonging to two groups. The directions of the first group characterize ten intrusions including the dated  $1970 \pm 3$  Ma Unoi sill and  $1976 \pm 9$  Ma Suna River Canyon dolerite, the corresponding paleomagnetic pole is  $44.4^\circ\text{N}$ ,  $101.5^\circ\text{E}$ ,  $A_{95} = 6.3^\circ$ . The second group comprises two intrusions including the  $1984 \pm 8$  Ma Pudozhgora intrusion and Krestoviy Navolok dyke with the corresponding paleopole calculated from 5 site mean poles is  $60.9^\circ\text{N}$ ,  $144.8^\circ\text{E}$ ,  $A_{95} = 6.8^\circ$ . Both remanence directions are supported by robust baked contact tests. We propose the first group's pole as the key 1975 Ma Fennoscandian pole. The second one is well dated, but based only on two intrusions without proper averaging of the paleosecular variations. We have also carried out a complimentary paleomagnetic study of the previously investigated 2504 Ma Shalskiy gabbro-norite dyke. The remanence of this dyke is now supported by the inverse contact test and statistics can be improved. Using our 1975 Ma pole together with coeval poles from Superior, Slave and Amazonia cratons we propose a provisional 1975 Ma paleogeographic reconstruction.

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## 1. Introduction

The 2.0–1.8 Ga time interval in Earth's history is marked by world-spread orogenic events (e.g. Zhao et al., 2009 and references therein). These events initiated assemblies of Laurentia (the cratonic part of North America), Baltica, Siberia, Australia, Kalahari cratons (e.g. Hoffman, 1988; Bogdanova et al., 2008; Rosen et al., 1994; Pisarevsky et al., 2008; Cawood and Korsch, 2008; Hanson et al., 2011) and accretionary growth of others (e.g. Brito Neves, 2011). Most of these and other cratons eventually collided to form the Paleo-Mesoproterozoic supercontinent Nuna (or Columbia) (Rogers and Santosh, 2002; Zhao et al., 2004, 2006; Meert, 2012 and references therein). The time of the final assembly of Nuna/Columbia is debated and estimated between ca. 1800 Ma (Zhang

et al., 2012 and references therein) and ca. 1600 Ma (Pisarevsky et al., 2014 and references therein).

Recent studies demonstrate that Baltica (the East European Craton) has not been assembled until ca. 1.8–1.7 Ga (Bogdanova et al., 2008; 2013; Lubnina et al., 2015a). At ca. 2.0 Ga, at least three proto-cratons – Kola, Karelia and Volgo-Sarmatia were separated by oceans (Melezhik and Hanski, 2013; Bogdanova et al., 2008, 2013). The small Kola Ocean between Kola and Karelia was closed at ca. 1.94 Ga (Daly et al., 2006; Melezhik and Hanski, 2013). Starting from ca. 1.90 Ga, the united Kola-Karelia grew in SW- to SSW direction (hereafter – in present day coordinates) by consequent accretion of island arcs and microcontinents to amalgamate Fennoscandia (Bogdanova et al., 2015).

Several early Paleoproterozoic paleogeographic reconstructions around 2.5–2.4 Ga (e.g. Bleeker and Ernst, 2006; Söderlund et al., 2010; Salminen et al., 2014; Pisarevsky et al., 2015) and late Paleoproterozoic, younger than 1.8 Ga (e.g. Pisarevsky and Bylund, 2010; Zhang et al., 2012; Pisarevsky et al., 2014) have been

\* Corresponding author.

E-mail address: [natalia.lubnina@gmail.com](mailto:natalia.lubnina@gmail.com) (N.V. Lubnina).